

University of Vienna

QED corrections for precision experiments

Yannick Ulrich

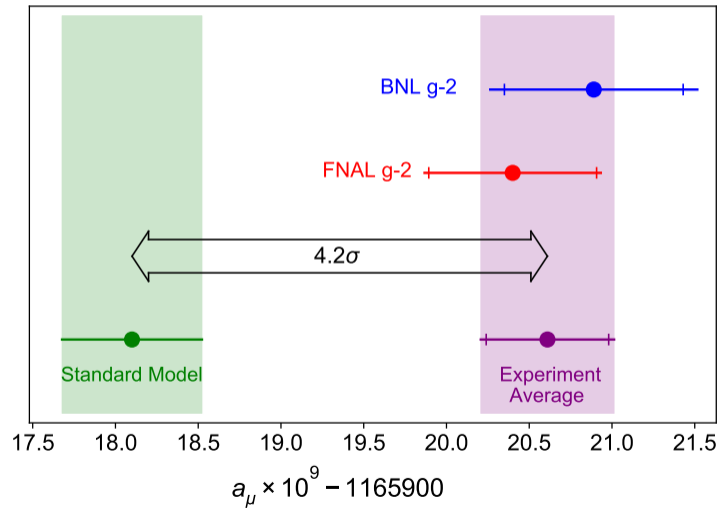
IPPP, University of Durham

18TH SEPTEMBER 2022

I hope to address the following

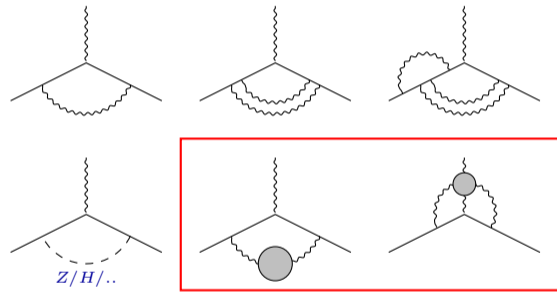
- $\alpha_{\text{QED}} \ll 1$, so why bother?
- ⇒ where do QED corrections matter?
- what challenges?
 - how to solve them (in pictures!)
 - some phenomenology (more pretty pictures!)
 - vision of the future

most precise measurement of $g - 2$



⇒ needs precise theory

many Feynman diagrams, incl. non-perturbative



theory uncertainty from hadronic physics

using optical theorem $s > 0$

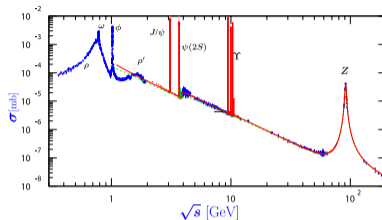
$$\int ds \left(K(s) \right) \left(\text{Diagram: two lines merging into a wavy line which then splits into two lines} \right)$$

⇒ very messy!

using $t < 0$

$$\int dt \left(K'(t) \right) \left(\text{Diagram: two lines merging into a circle which then splits into two lines} \right)$$

⇒ much cleaner but **smaller**



target accuracy: 10^{-5} (\rightarrow 1% on HVP)

- dominant NNLO corr. with full m dep.

[Carlson Calame et al. 20; Banerjee, Engel, Signer, YU 20]

- full NNLO corr. (currently w/o m^2/Q^2 , add later) [Broggio, Engel, Ferroglia, Mandal, Mastrolia, Passera, Rocco, Ronca, Signer, Torres Bobadilla, Zoller, YU 2?]

- electronic N³LO w/o m^2/Q^2
- resummation

$$\begin{aligned}
 \sigma = & \int d\Phi_2 \left| \begin{array}{c} \text{tree} \\ \text{1-loop} \\ \text{2-loop} \\ \text{3-loop} \\ \dots \end{array} \right|^2 \\
 & + \int d\Phi_3 \left| \begin{array}{c} \text{1-loop} \\ \text{2-loop} \\ \text{3-loop} \\ \dots \end{array} \right|^2 \\
 & + \int d\Phi_4 \left| \begin{array}{c} \text{2-loop} \\ \text{3-loop} \\ \dots \end{array} \right|^2 \\
 & + \int d\Phi_5 \left| \begin{array}{c} \text{3-loop} \\ \dots \end{array} \right|^2 \\
 & + \text{LL} + \text{NLL} + \dots
 \end{aligned}$$

the world is not just $g - 2...$

- luminosity measurements $\Rightarrow e^+e^- \rightarrow e^+e^-$ (Belle, FCC-ee, ...)
[Banerjee, Engel, Schalch, Signer, YU 21]
- dark sector searches $\Rightarrow e^+e^- \rightarrow \gamma\gamma$ (PADME, also for luminosity...)
[Engel, Naterop, Signer, YU, Zoller 2?]
- R ratios $\Rightarrow e^+e^- \rightarrow \mu^+\mu^-$ (DAΦNE, VEPP, ...)
- τ physics $\Rightarrow e^+e^- \rightarrow \tau^+\tau^-$ (Belle) [Kollatzsch, YU 2?]
- proton radius $\Rightarrow lp \rightarrow lp$ and $ee \rightarrow ee$ (P2, PRad, MUSE)
[Bucoveanu, Spiesberger 18; Banerjee, Engel, Signer, YU 20; Banerjee, Engel, Schalch, Signer, YU 21]
- lepton decays $\Rightarrow l \rightarrow l'\nu\bar{\nu} + \{ee, \gamma, \gamma\gamma\}$ (MEG, Mu3e, Belle, ...)
[Pruna, Signer, YU 16; YU, 17; Engel, Gnendiger, Signer, YU, 18, Banerjee, Coutinho, Engel, Gurgone, Signer, YU 2?]



MCMULE

a framework for QED corrections

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basic strategy: use 40+ years of QCD experience on QED

- use automation where available and useful (eg. OpenLoops [Buccioni, Pozzorini, Zoller 18; Buccioni, Lang, Lindert, Maierhöfer, Pozzorini 19])
- adapt QCD results where known (eg. [Bernreuther et al., 04])
- use methods invented (eg. [Frixione, Kunstz, Signer 96])

QED and QCD calculations have many common issues, but . . .

- Abelian structure \Rightarrow a bit easier [no big deal]
- much simpler infrared structure [advantage]
- want/need $m \neq 0$ since $\log m$ physical [problem]
- more exclusive, e.g. hard collinear emission [problem]

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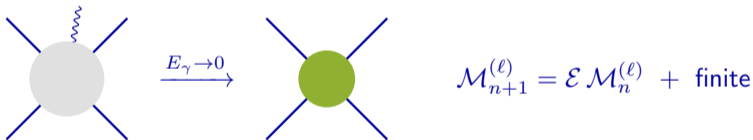
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soft singularities

$$\int d\Phi_\gamma \quad \text{diagram} \sim \int_0 dE_\gamma E_\gamma \int_{-1}^1 d(\cos\theta) \frac{1}{E_\gamma^2(1 - \beta \cos\theta)}$$

⇒ luckily universality of soft singularities



for any process and loop order. Similarly for virtual

$$e^{\hat{\epsilon}} \sum_{\ell=0}^{\infty} \mathcal{M}_n^{(\ell)} = \text{finite}$$

⇒ subtraction scheme at any order (FKS^ℓ) [Engel, Signer, YU 19]

$$\underbrace{\int d\Phi_\gamma}_{\text{divergent and complicated}} \underbrace{\text{diagram}}_{\text{divergent and complicated}} = \underbrace{\int d\Phi_\gamma}_{\text{complicated but finite}} \left(\underbrace{\text{diagram}}_{\text{divergent and complicated}} - \underbrace{\text{diagram}}_{\text{finite}} \right) + \underbrace{\int d\Phi_\gamma}_{\text{divergent but easy}} \underbrace{\text{diagram}}_{\text{finite}}$$

- very QCD-y
- based on [Frixione, Kunszt, Signer 96]
- no resolution parameter or photon mass, just DREG
- unphysical $0 < \xi_c \lesssim 1$ to test stability, implementation, ...

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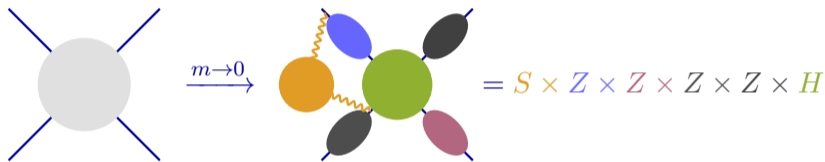
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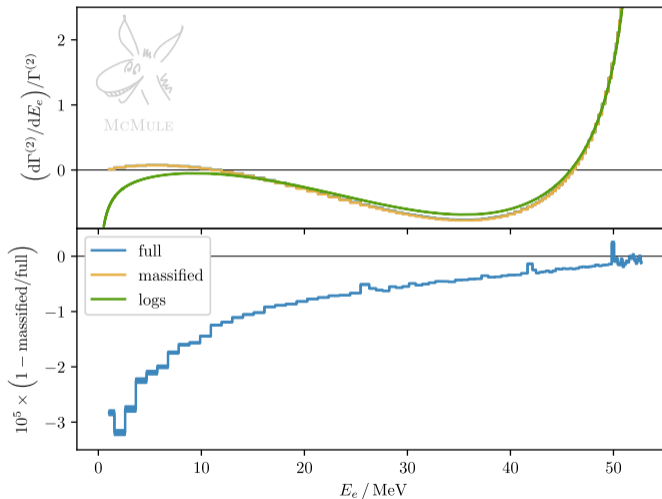
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- loop integrals with internal masses are very complicated!
- **but** $m_e^2 \ll m_\mu^2 \sim Q^2$ for many applications
- ⇒ don't actually care about full m_e dependence
- **but** $\int \langle \text{expanded integrand} \rangle \neq \langle \text{expanded integral} \rangle$
- ⇒ method of regions [Beneke, Smirnov 98] (**hard**, **soft**, **collinear**, ...)

universality of collinear singularities \rightarrow calculate up to $\mathcal{O}(m^2/Q^2)$



- H : hard function $\sim \left. \begin{array}{c} \diagup \quad \diagdown \\ \text{grey circle} \\ \diagdown \quad \diagup \end{array} \right|_{m=0}$
- Z : process independent jet function
- S : simple soft function



[Chen 18] v. [Engel, Gnendiger, Signer, YU 18] v. [Arbuzov et al. 02]

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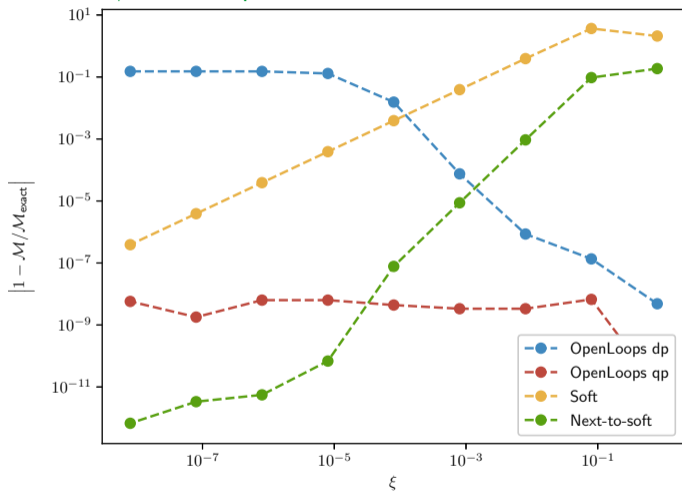
$$\mathcal{M}_{n+1}^{(\ell)} \sim \frac{1}{E_\gamma^2(1-\beta \cos \theta)}$$

real-virtual (or even real-real-virtual)

- ‘trivial’ in principle [Buccioni, Pozzorini, Zoller 18; Buccioni, Lang, Lindert, Maierhöfer, Pozzorini et al. 19]
 - extremely delicate numerically for $E_\gamma \rightarrow 0$ (or $\cos \theta \rightarrow 1$)
- ⇒ Taylor expand around $E_\gamma = 0$ if small
- LBK theorem [Low 58; Burnett, Kroll 67] and extension [Engel, Signer, YU 21]



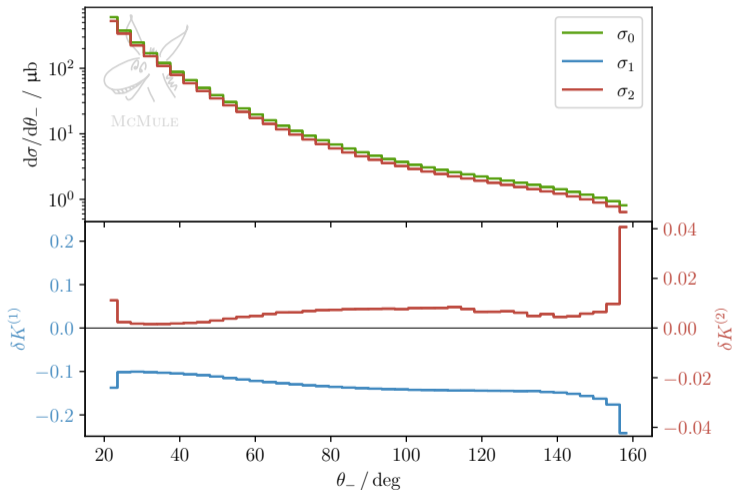
example $e^+e^- \rightarrow e^+e^-\gamma$ @ one-loop



compare with exact calculation in Mathematica

[Banerjee, Engel, Schalch, Signer, YU 21]

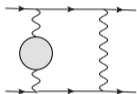
$\sqrt{s} = 1020 \text{ MeV}$



$E_{\pm} > 408 \text{ MeV}, 20^{\circ} \leq \theta_{\pm} \leq 160^{\circ}, |180^{\circ} - \theta_+ - \theta_-| < 10^{\circ}$

a few more hurdles

- VP diagrams for $e/\mu/\tau/\text{had}/\dots$ numerically with full mass dependence



- collinear pseudo-singularities $\lim_{\rightarrow 0} \sphericalangle(p_\gamma, p_i) \Rightarrow L$
- phase-space tuning s.t. $\cos \sphericalangle \sim x_i$

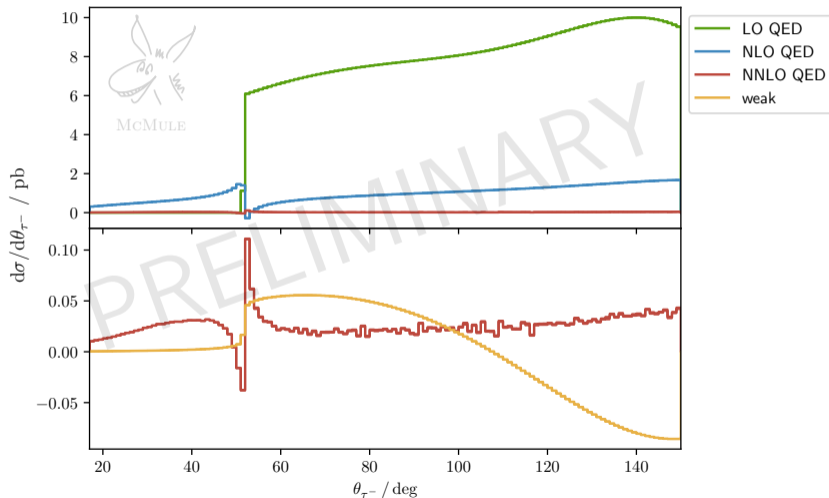
\Rightarrow at most one small angle \rightarrow FKS partitioning

- polarisation & EW where applicable

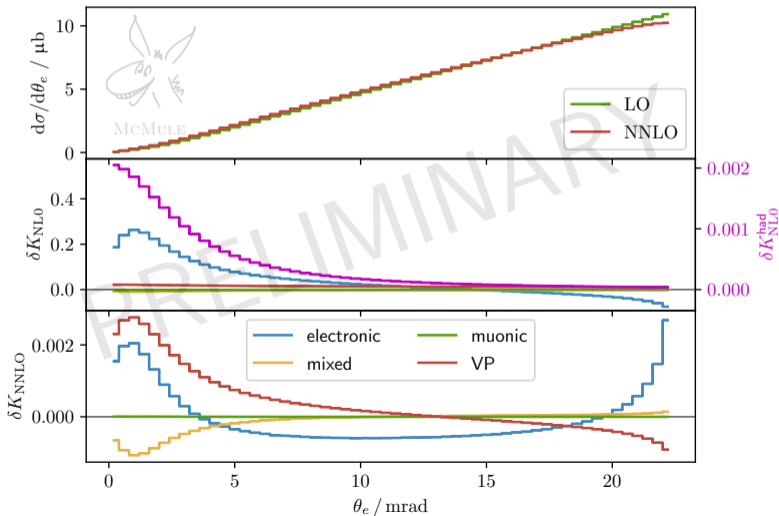


[Signer 22]

for Belle II, polarised initial state, NLO-EW \oplus dominant NNLO [Kollatzsch, YU 2?]



$E_{\text{beam}}^\mu = 150 \text{ GeV}$, $E_e > 1 \text{ GeV}$, $\theta_\mu > 0.3 \text{ mrad}$ [Broggio, Engel, Ferroglia, Mandal, Mastrolia, Passera, Rocco, Ronca, Signer, Torres Bobadilla, Zoller, YU 2?]

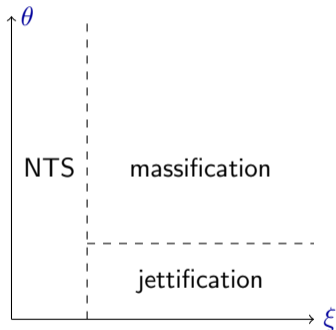


$ee \rightarrow \gamma^*$ can be taken to N^3 LO

- VVV: known
[Fael, Lange, Schönwald, Steinhauser 22]
 - RRR: “trivial”
 - RRV: OpenLoops + NTS stabilisation
 - RVV: massless known (three-jet @ NNLO), massive (DiffExp?)
- ⇒ LBK + jettification at two-loop

jettification

- expand for small emission angles



the NNLO era is here, not only for QCD, also for QED

future steps

- more NNLO QED \oplus EW
 - NNLO QED \oplus PS
 - higher energies
 - massification for real corrections
 - collinear stabilisation
- N³ LO for $\gamma^* \rightarrow ll$
 \Rightarrow Workstop in Durham





McMULE

mule-tools.gitlab.io

f.l.t.r.: F.Hagelstein (Mainz), A.Coutinho (IFIC Valencia), N.Schalch (Bern), L.Naterop (Zurich & PSI), S.Kollatzsch (Zurich & PSI), A.Signer (Zurich & PSI), M.Rocco (PSI), T.Engel (→ Freiburg), V.Sharkovska (Zurich & PSI), Y.Ulrich (Durham), A.Gurgone (Pavia), not shown: P. Banerjee (Zhejiang), A. Proust (Lyon)

$$\begin{aligned}
 & \text{Diagram with wavy line} \xrightarrow{E_\gamma \rightarrow 0} \underbrace{\frac{1}{E_\gamma^2} \text{Diagram with green blob}}_{\text{universal eikonal}} + \underbrace{\frac{1}{E_\gamma} \text{Diagram with orange and green blobs}}_{\text{next-to-soft}} + \mathcal{O}(E_\gamma^0) \\
 & = \mathcal{E} \times \text{Diagram with grey blob} + D_{\text{LBK}} \left[\text{Diagram with grey blob} \right] + \left(\sum_{ijk} \mathcal{S}_{ijk} \right) \times \text{Diagram with grey blob}
 \end{aligned}$$